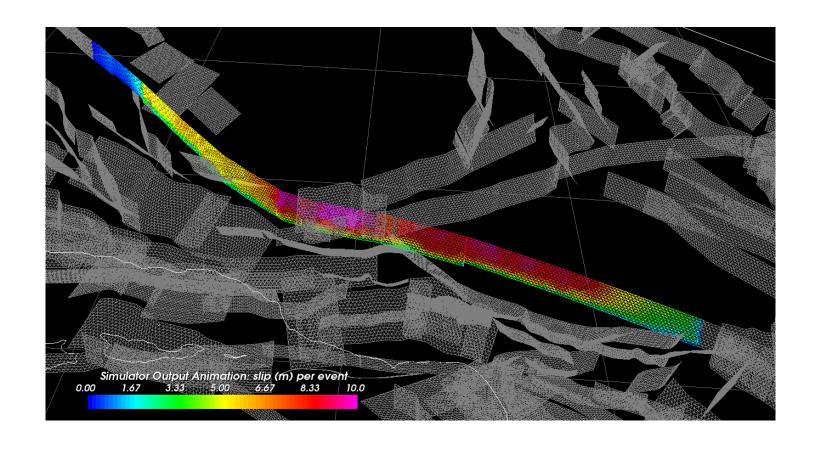


Data Integration and Visualization Tools for Bringing Paleoseismic Data and Simulator (& UCERF3) Results Together



Kevin R. Milner

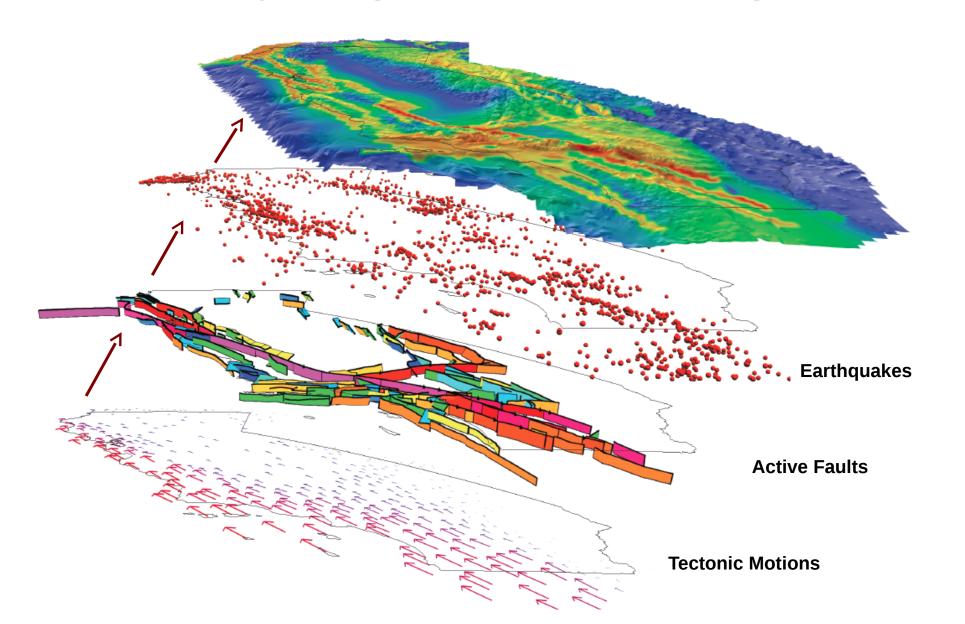
University of Southern California

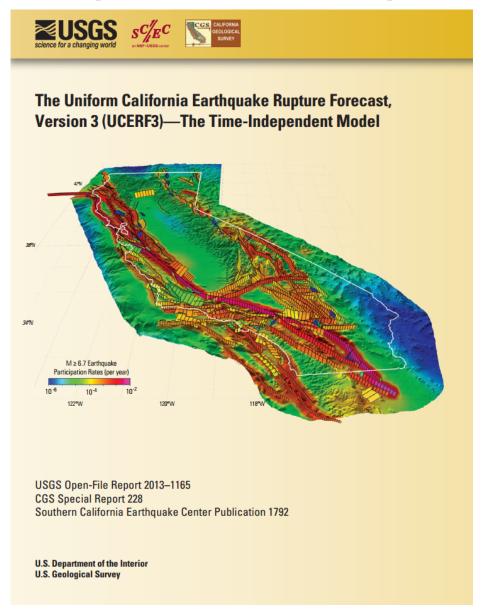
SoSAFE Workshop, 2016 SCEC Annual Meeting September 10, 2016



UCERF3

Uniform California Earthquake Rupture Forecast (UCERF3) by the Working Group on California Earthquake Probabilities (Field et al., 2014)







Paleoseismic Data and Slip Rates Battle It Out in UCERF3

UCERF3 "Grand Inversion"

- Solves for event rates of fault-based ruptures
- Data constraints include:
 - Paleoseismic data
 - Slip rates
 - Regional Magnitude Frequency Distributions
 - Fault Section Magnitude Frequency Distributions
- Paleoseismic data included:
 - 31 Event Rate Constraints
 - 23 Mean Slip Constraints
- The problem: can't perfectly fit paleo & slip rate data simultaneously
 - Final UCERF3 model constraint weights represent a compromise between the two datasets

Table 5. The grand inversion system of equations used in solving for the long-term rate of fault-based ruptures, where f_r represents the frequency or rate of the rth rupture.

[Important implementation details, such as equation-set weighting, are given in appendix N (this report)]

portant implementation details, such a		s equation-set weighting, are given in appendix iv (this report)]
Equation set		Description
$\sum_{r=1}^{R} D_{sr} f_r = v_s$	(1)	Slip Rate Balancing: v_s is the subsection slip rate (from a deformation model) and D_{sr} is the slip on the s^{th} subsection in the r^{th} event, averaged over multiple occurrences of the ruptur and as measured at mid-seismogenic depth.
$\sum_{r=1}^{R} G_{sr} P_r^{paleo} f_r = f_s^{paleo}$	(2)	Paleoseismic Event Rate Matching: f_s^{paleo} is a paleoseismically inferred event rate estimate, G_{sr} specifies whether the \mathbf{r}^{th} rupture utilizes the \mathbf{s}^{th} subsection (0 or 1), and P_r^{paleo} is the probability that the \mathbf{r}^{th} rupture would be seen in a paleoseismic trench.
$R_s^m = \frac{R_{s-1}^m + R_{s+1}^m}{2}$	(3)	Fault Section Smoothness Constraint : This enables forcing the nucleation rate, R, in the m th magnitude bin to vary smoothly along a fault section, where the s-1 and s+1 subsections are adjacent to the s th subsection.
$\lambda_r f_r = 0$	(4)	Improbability Constraint : This allows us to force relatively improbable events to have a lower rate (for example, based on multi-fault rupture likelihoods). A higher value of adds more misfit for a given rupture rate, forcing the inversion to minimize that rupture rate further.
$f_r = f_r^{a-priori}$	(5)	A Priori Constraint : This constrains the rates of particular ruptures to target values, either on an individual basis (for example, make Parkfield occur every ~25 years) or for a complete rupture set (for example, as close as possible to thos in UCERF2).
$\sum_{r=1}^{R} M_{gr}^{m} f_{r} = R_{g}^{m}$	(6)	Regional MFD Constraint: This enables forcing a geographi region, g, to have a specified magnitude-frequency distribution (MFD), such as Gutenberg-Richter. R_g^m represents the nucleation rate for the m^{th} magnitude bin in the g^{th} region. Matrix $M_{gr}^m f_r$ contains the product of whether the r^{th} rupture falls in the m^{th} magnitude bin (0 or 1) multiplied by the fraction of that rupture that nucleates in the g^{th} region.
$\sum_{r}^{R} M_{sr}^{m} f_{r} = R_{s}^{m}$	(7)	Fault Section MFD Constraint: This enables forcing subsections to have specific nucleation MFDs. $R_{s_b}^m$ is the

nucleation rate for the mth magnitude bin on the sth subsection.

Matrix $M_{sr}^{m}f_{r}$ contains the product of whether the r^{th} rupture falls in the m^{th} magnitude bin (0 or 1) multiplied by the fraction

of that rupture that nucleates on the sth subsection.



Paleoseismic Data and Slip Rates **Battle It Out in UCERF3**

---- Model

Data

Model

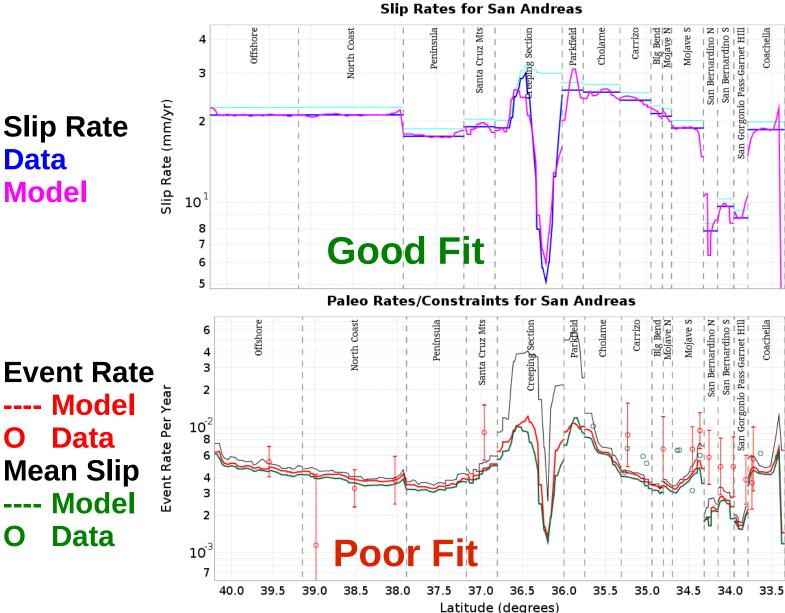
Data

High Paleoseismic Constraint Weight

Slip Rates for San Andreas

Slip Rate (mm/yr) Slip Rate **Data** Model **Poor Fit** Paleo Rates/Constraints for San Andreas 10-2 Event Rate Per Year **10**⁻³ **Mean Slip** Good Fit Latitude (degrees)

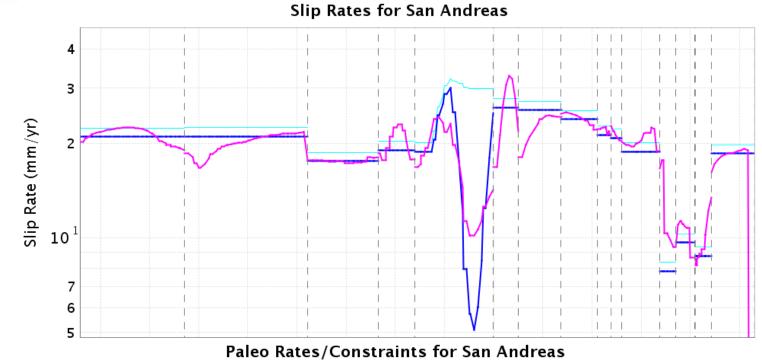
High Slip Rate Constraint Weight



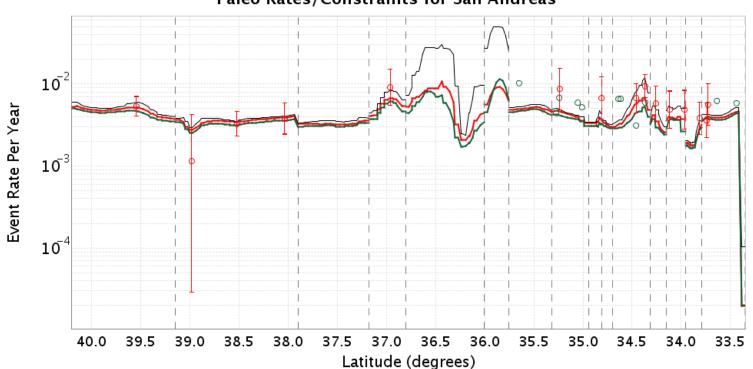


UCERF3 Final Model Fits





Event Rate
---- Model
O Data
Mean Slip
---- Model
O Data

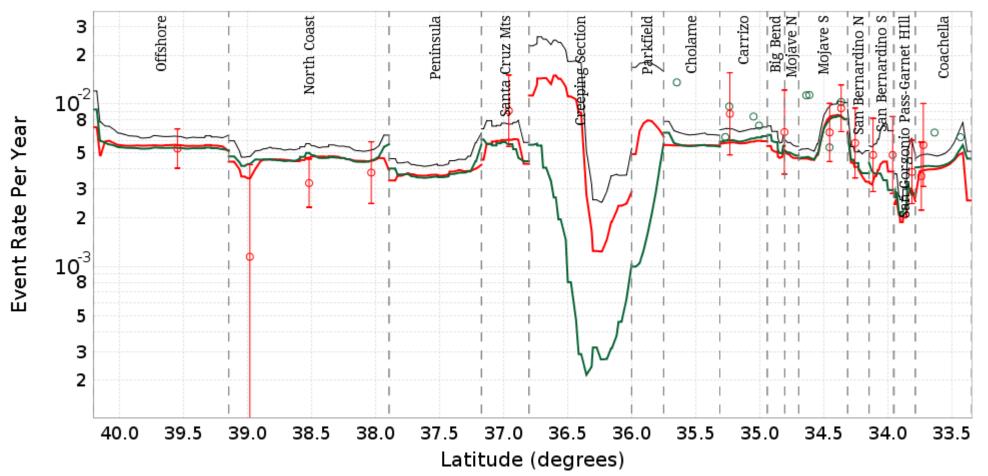




RSQSim to the Rescue?

- RSQSim fits slip rates exactly (backslip model)
- Paleoseismic fits are similar to UCERF3 on the SAF
- Magnitude Frequency Distributions don't match expected (yet)

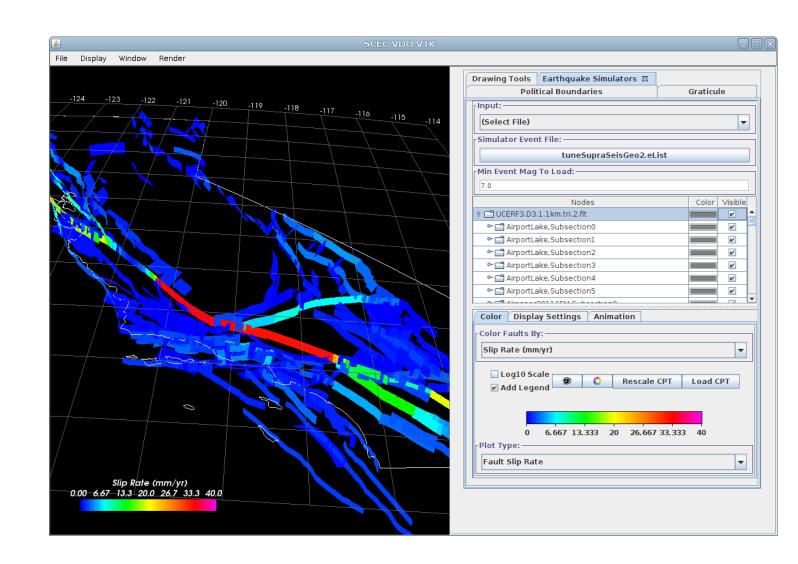






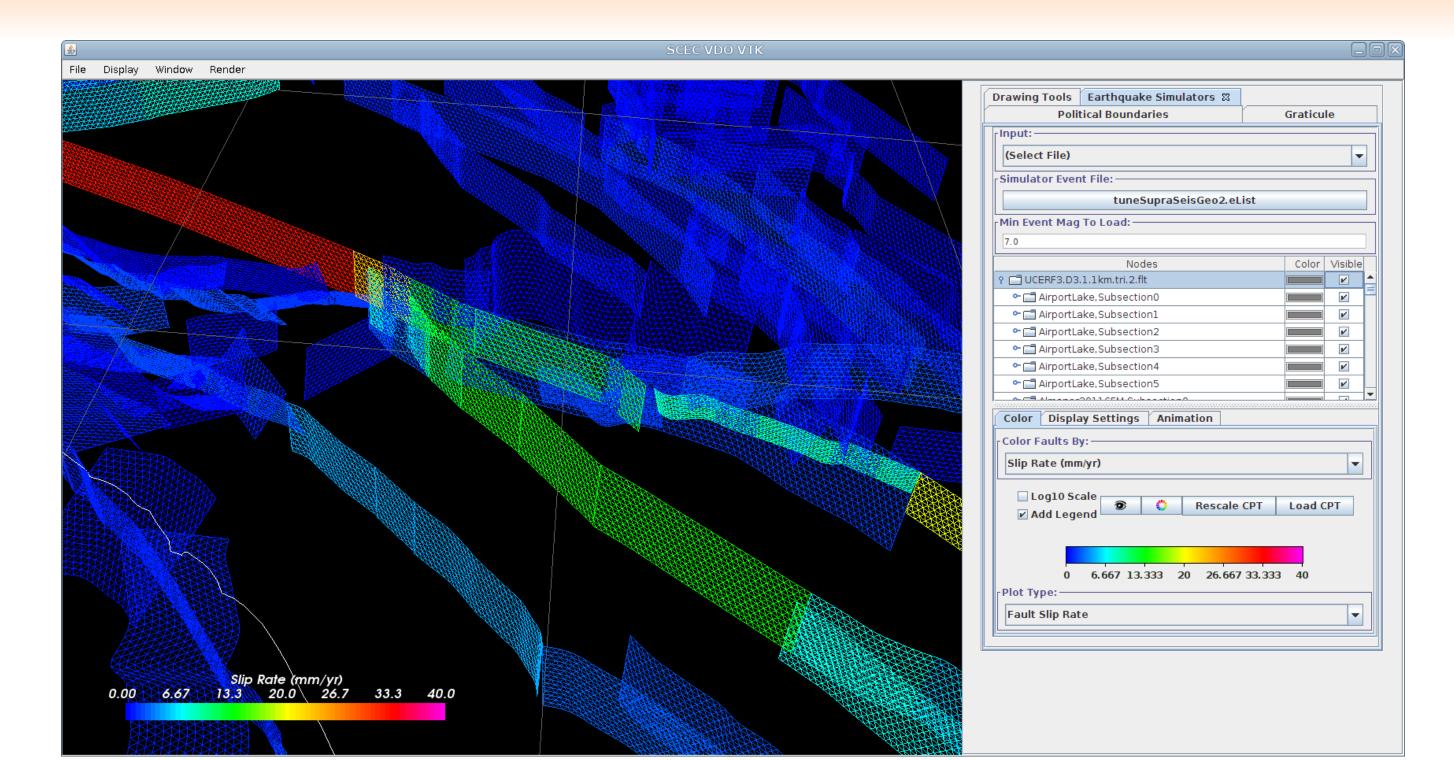
SCEC-VDO

- Rebuilt from the ground up in 2016
 - Replaces outdated Java3D library
 - Runs on Windows, Mac OS X, Linux
- Download: scedvdo.usc.edu
- 2016 SCEC UseIT Internship Focused on RSQSim
 - New exploration and movie making tools available
 - Stop by during poster sessions for a demo



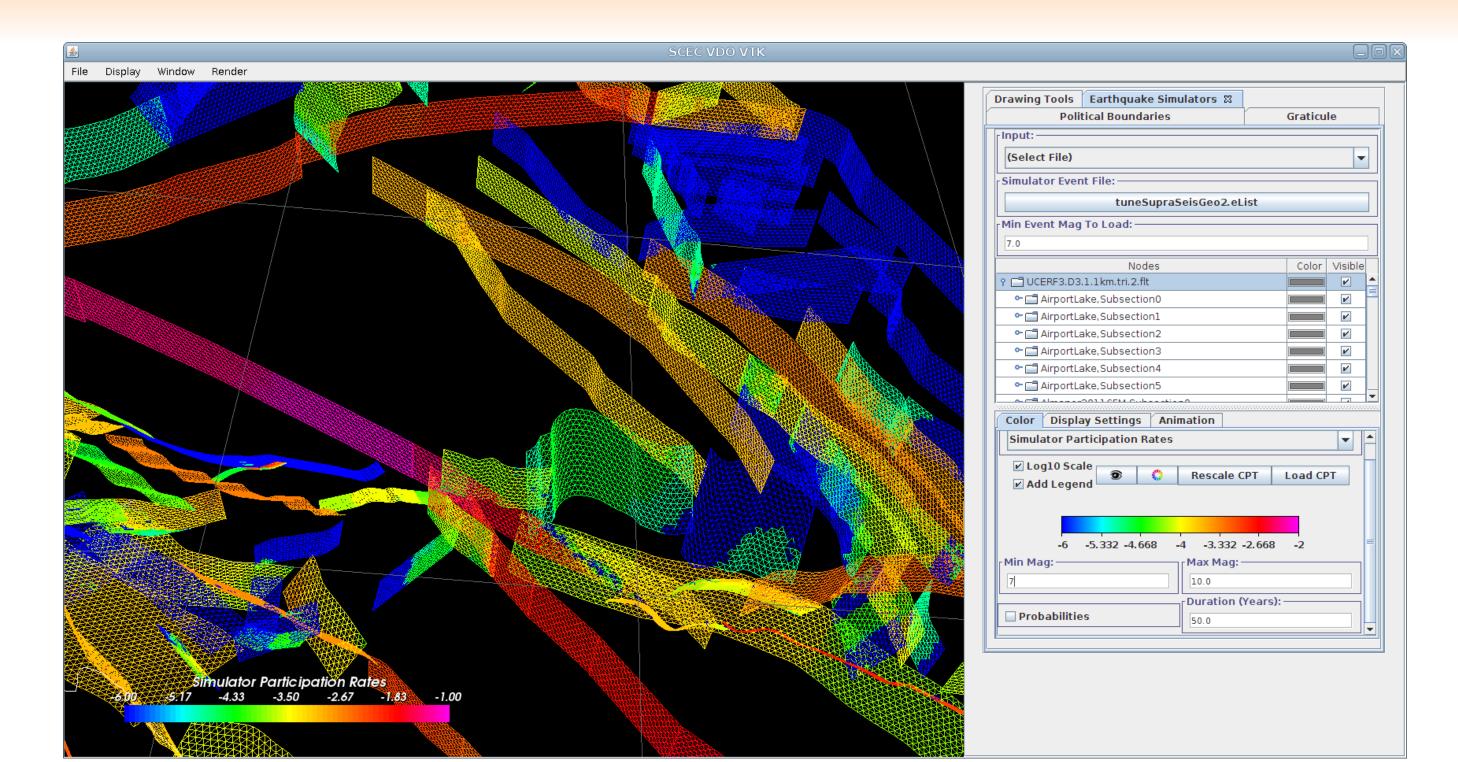


SCEC-VDO: RSQSim Slip Rates





SCEC-VDO: RSQSim Participation Rates

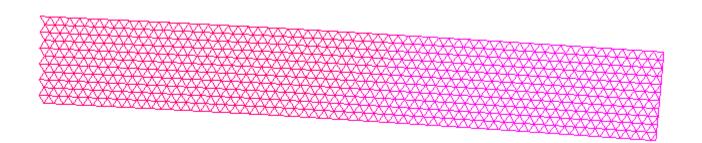




SCEC-VDO: RSQSim MFDs

Resluts can be viewed at 3 scales

 Fault Section: UCERF3 fault section, e.g. SAF Mojave S



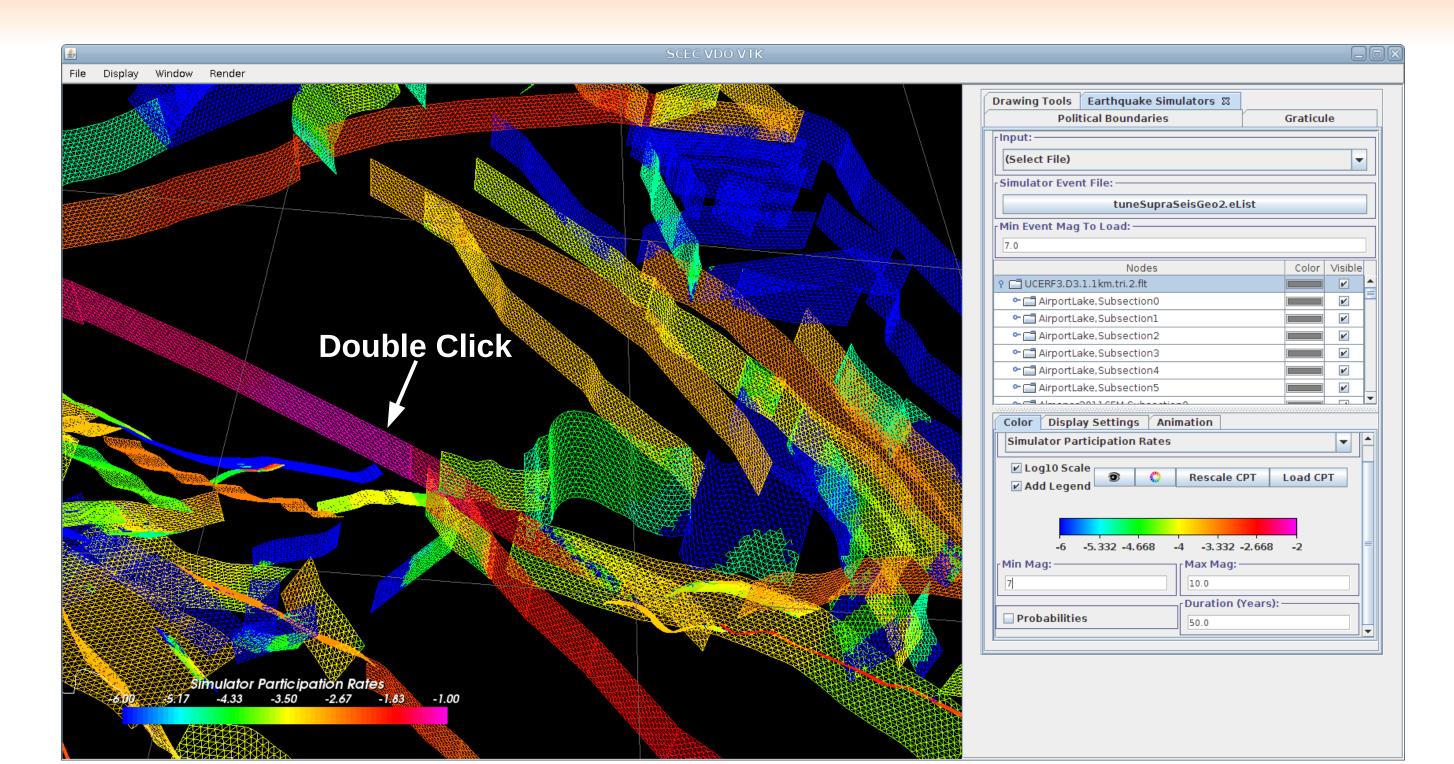
• Fault Subsection: Smallest UCERF3 unit where length ~= ½ DDW



• Patch: a single triangle with ~1km² area

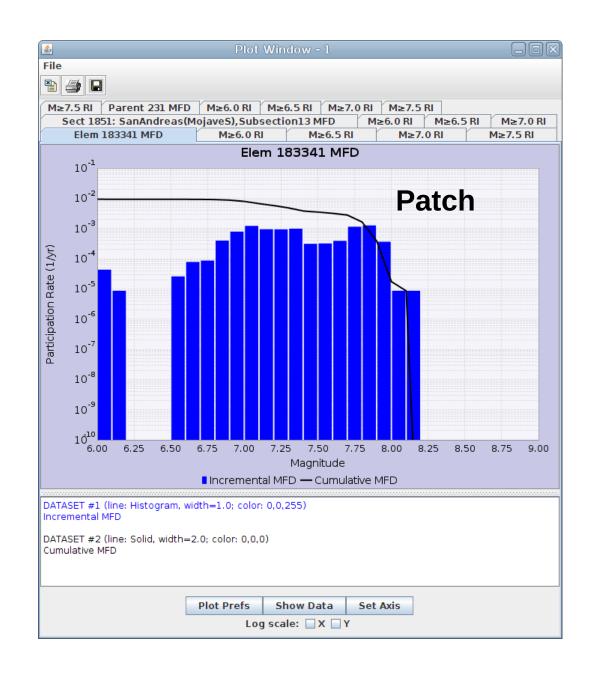


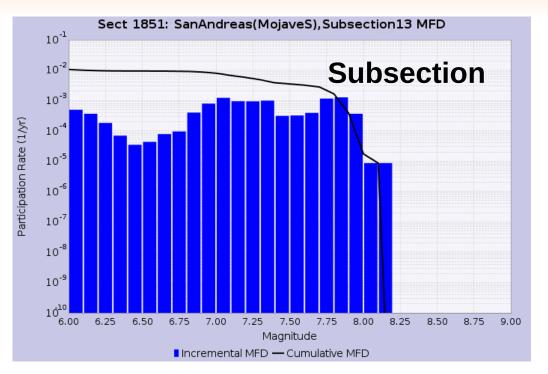
SCEC-VDO: RSQSim MFDs

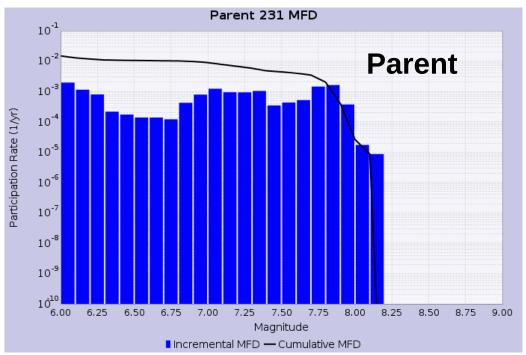




SCEC-VDO: RSQSim MFDs



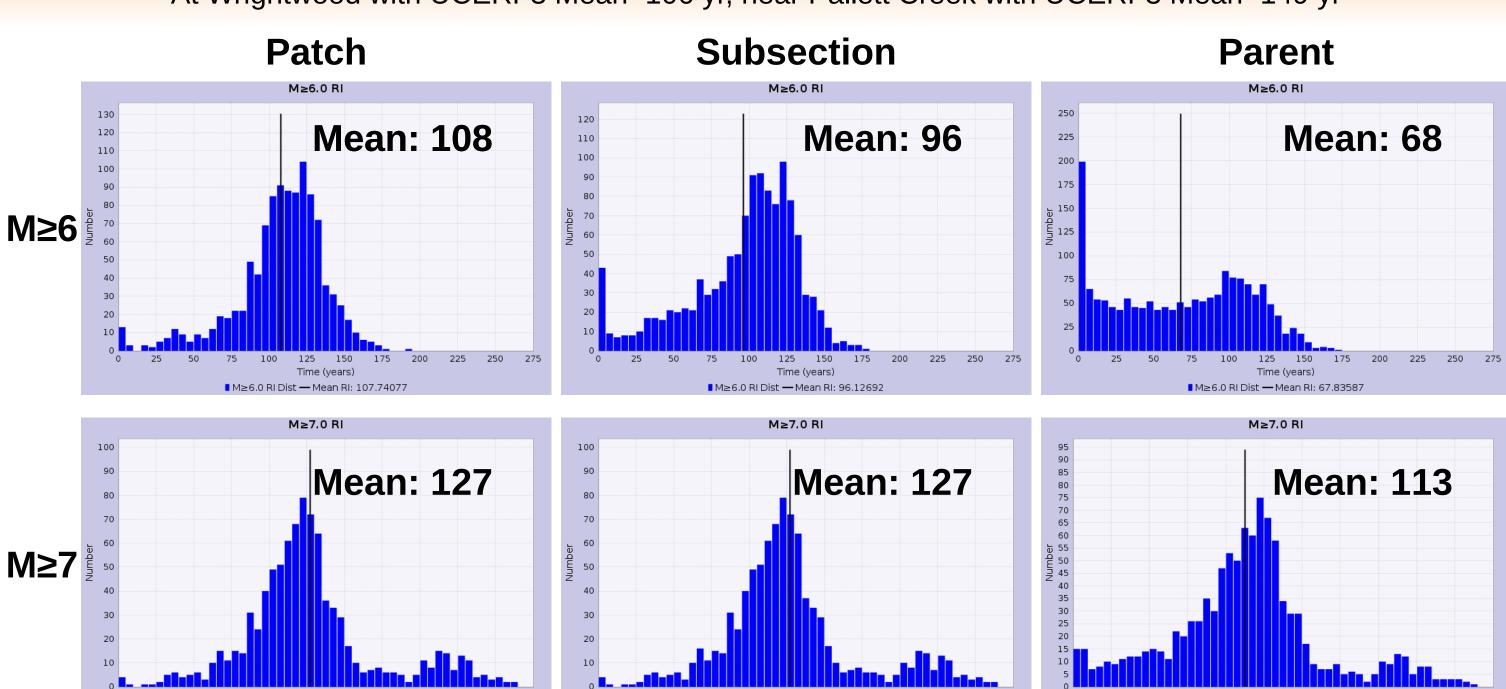






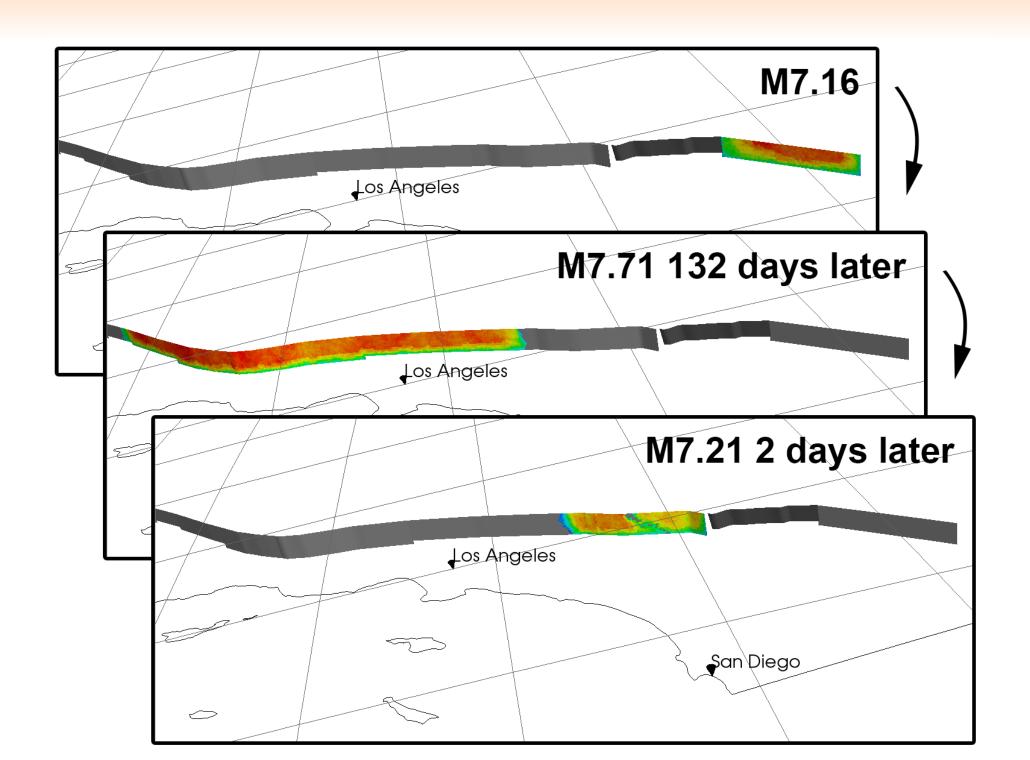
SCEC-VDO: RSQSim Recurrence Intervals

At Wrightwood with UCERF3 Mean=106 yr, near Pallett Creek with UCERF3 Mean=149 yr





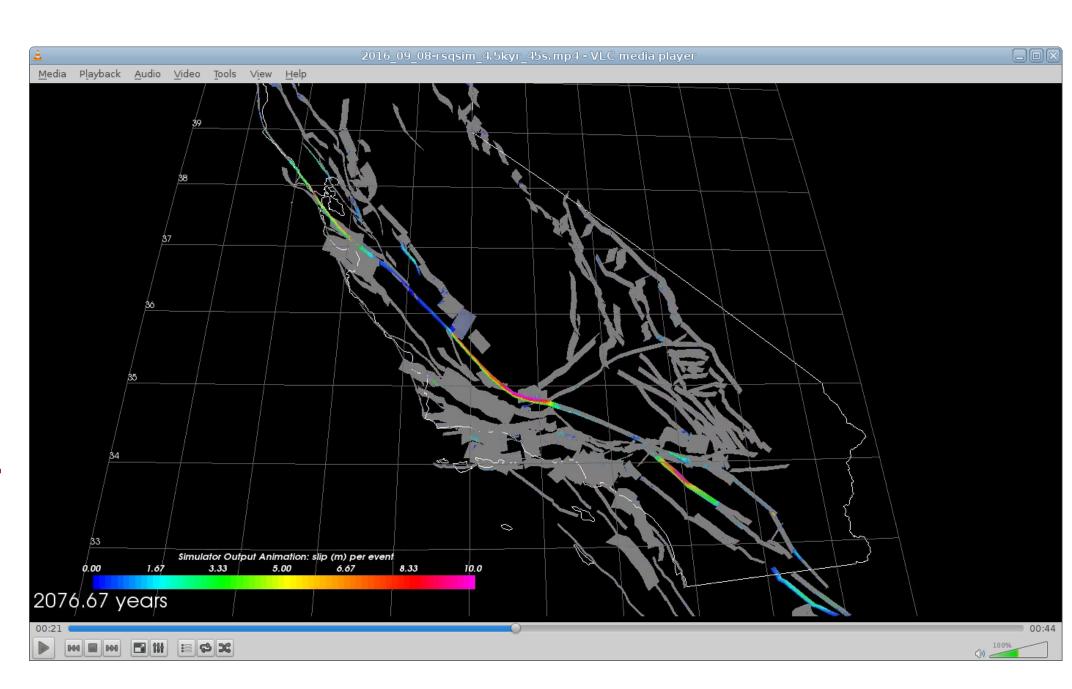
SCEC-VDO: View Sequences





SCEC-VDO: Render Movies

- Allows for visual inspection of patterns
- Previous events fade out over time
 - 50 year fade works well
- Can visualize thousands of years or specific sequences





Thank you!